

MOMENTS OF INERTIA OF MASSES

9.11 MOMENT OF INERTIA OF A MASS

Consider a small mass Δm mounted on a rod of negligible mass which can rotate freely about an axis AA' (Fig. 9.20*a*). If a couple is applied to the system, the rod and mass, assumed to be initially at rest, will start rotating about AA'. The details of this motion will be studied later in dynamics. At present, we wish only to indicate that the time required for the system to reach a given speed of rotation is proportional to the mass Δm and to the square of the distance *r*. The product $r^2 \Delta m$ provides, therefore, a measure of the *inertia* of the system, i.e., a measure of the resistance the system offers when we try to set it in motion. For this reason, the product $r^2 \Delta m$ is called the *moment of inertia* of the mass Δm with respect to the axis AA'.



Consider now a body of mass m which is to be rotated about an axis AA' (Fig. 9.20b). Dividing the body into elements of mass Δm_1 , Δm_2 , etc., we find that the body's resistance to being rotated is measured by the sum $r_1^2 \Delta m_1 + r_2^2 \Delta m_2 + \cdots$. This sum defines, therefore, the moment of inertia of the body with respect to the axis AA'. Increasing the number of elements, we find that the moment of inertia is equal, in the limit, to the integral

$$I = \int r^2 \, dm \tag{9.28}$$

The radius of gyration k of the body with respect to the axis AA' is defined by the relation

$$I = k^2 m$$
 or $k = \sqrt{\frac{I}{m}}$ (9.29)

The radius of gyration k represents, therefore, the distance at which the entire mass of the body should be concentrated if its moment of inertia with respect to AA' is to remain unchanged (Fig. 9.20c). Whether it is kept in its original shape (Fig. 9.20b) or whether it is concentrated as shown in Fig. 9.20c, the mass m will react in the same way to a rotation, or gyration, about AA'.

If SI units are used, the radius of gyration k is expressed in meters and the mass m in kilograms, and thus the unit used for the moment of inertia of a mass is kg \cdot m². If U.S. customary units are used, the radius of gyration is expressed in feet and the mass in slugs (i.e., in $lb \cdot s^2/ft$), and thus the derived unit used for the moment of inertia of a mass is $lb \cdot ft \cdot s^2$.[†]

The moment of inertia of a body with respect to a coordinate axis can easily be expressed in terms of the coordinates x, y, zof the element of mass dm (Fig. 9.21). Noting, for example, that the square of the distance r from the element dm to the y axis is $z^{2} + x^{2}$, we express the moment of inertia of the body with respect to the *y* axis as

$$I_y = \int r^2 dm = \int (z^2 + x^2) dm$$

Similar expressions can be obtained for the moments of inertia with respect to the x and z axes. We write

$$I_x = \int (y^2 + z^2) dm$$
$$I_y = \int (z^2 + x^2) dm$$
$$I_z = \int (x^2 + y^2) dm$$

(9.30)

Fig. 9.21

†It should be kept in mind when converting the moment of inertia of a mass from U.S. customary units to SI units that the base unit *pound* used in the derived unit $lb \cdot ft \cdot s^2$ is a unit of force (not of mass) and should therefore be converted into newtons. We have

1 lb
$$\cdot$$
 ft \cdot s² = (4.45 N)(0.3048 m)(1 s)² = 1.356 N \cdot m \cdot s²

or, since $1 \text{ N} = 1 \text{ kg} \cdot \text{m/s}^2$,

$$1 \text{ lb} \cdot \text{ft} \cdot \text{s}^2 = 1.356 \text{ kg} \cdot \text{m}^2$$







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9.12 PARALLEL-AXIS THEOREM

Consider a body of mass *m*. Let *Oxyz* be a system of rectangular coordinates whose origin is at the arbitrary point O, and Gx'y'z' a system of parallel centroidal axes, i.e., a system whose origin is at the center of gravity G of the body[†] and whose axes x', y', z' are parallel to the x, y, and z axes, respectively (Fig. 9.22). Denoting by $\overline{x}, \overline{y}, \overline{z}$ the coordinates of G with respect to Oxyz, we write the following relations between the coordinates x, y, z of the element dm with respect to Oxyz and its coordinates x', y', z' with respect to the centroidal axes Gx'y'z':

$$x = x' + \overline{x} \qquad y = y' + \overline{y} \qquad z = z' + \overline{z} \qquad (9.31)$$

Referring to Eqs. (9.30), we can express the moment of inertia of the body with respect to the *x* axis as follows:

$$\begin{split} I_x &= \int (y^2 + z^2) \, dm = \int \left[(y' + \overline{y})^2 + (z' + \overline{z})^2 \right] \, dm \\ &= \int (y'^2 + z'^2) \, dm + 2\overline{y} \int y' \, dm + 2\overline{z} \int z' \, dm + (\overline{y}^2 + \overline{z}^2) \int dm \end{split}$$

The first integral in this expression represents the moment of inertia $I_{x'}$ of the body with respect to the centroidal axis x'; the second and third integrals represent the first moment of the body with respect to the z'x' and x'y' planes, respectively, and, since both planes contain G, the two integrals are zero; the last integral is equal to the total mass m of the body. We write, therefore,

$$I_x = \bar{I}_{x'} + m(\bar{y}^2 + \bar{z}^2)$$
(9.32)

and, similarly,

$$I_{y} = \bar{I}_{y'} + m(\bar{z}^{2} + \bar{x}^{2}) \qquad I_{z} = \bar{I}_{z'} + m(\bar{x}^{2} + \bar{y}^{2})$$
(9.32')

We easily verify from Fig. 9.22 that the sum $\overline{z}^2 + \overline{x}^2$ represents the square of the distance OB between the y and y' axes. Similarly, $\overline{y}^2 + \overline{z}^2$ and $\overline{x}^2 + \overline{y}^2$ represent the squares of the distance between the x and x' axes and the z and z' axes, respectively. Denoting by dthe distance between an arbitrary axis AA' and a parallel centroidal axis BB' (Fig. 9.23), we can, therefore, write the following general relation between the moment of inertia I of the body with respect to AA' and its moment of inertia I with respect to BB':

$$I = \overline{I} + md^2 \tag{9.33}$$

Expressing the moments of inertia in terms of the corresponding radii of gyration, we can also write

$$c^2 = \overline{k}^2 + d^2 \tag{9.34}$$

where k and k represent the radii of gyration of the body about AA'and BB', respectively.

[†]Note that the term *centroidal* is used here to define an axis passing through the center of gravity G of the body, whether or not G coincides with the centroid of the volume of the body.



Fig. 9.23

Slender rod	y G z L x	$I_y = I_z = \frac{1}{12}mL^2$
Thin rectangular plate	y G Z z	$\begin{split} I_x &= \frac{1}{12} m (b^2 + c^2) \\ I_y &= \frac{1}{12} m c^2 \\ I_z &= \frac{1}{12} m b^2 \end{split}$
Rectangular prism		$\begin{split} I_x &= \frac{1}{12} m(b^2 + c^2) \\ I_y &= \frac{1}{12} m(c^2 + a^2) \\ I_z &= \frac{1}{12} m(a^2 + b^2) \end{split}$
Thin disk	y z x	$\begin{split} I_x &= \frac{1}{2}mr^2\\ I_y &= I_z = \frac{1}{4}mr^2 \end{split}$
Circular cylinder	y z z	$\begin{split} I_x &= \frac{1}{2} m a^2 \\ I_y &= I_z = \frac{1}{12} m (3 a^2 + L^2) \end{split}$
Circular cone	z h r x	$\begin{split} &I_x = \frac{3}{10}ma^2 \\ &I_y = I_z = \frac{3}{5}m(\frac{1}{4}a^2 + h^2) \end{split}$
Sphere		$I_x = I_y = I_z = \frac{2}{5}ma^2$

Fig. 9.28 Mass moments of inertia of common geometric shapes.



SAMPLE PROBLEM 9.9

Determine the moment of inertia of a slender rod of length L and mass m with respect to an axis which is perpendicular to the rod and passes through one end of the rod.

SOLUTION

Choosing the differential element of mass shown, we write

$$dm = \frac{m}{L}dx$$

$$I_y = \int x^2 dm = \int_0^L x^2 \frac{m}{L} dx = \left[\frac{m}{L}\frac{x^3}{3}\right]_0^L \qquad I_y = \frac{1}{3}mL^2 \quad \checkmark$$





SAMPLE PROBLEM 9.10

For the homogeneous rectangular prism shown, determine the moment of inertia with respect to the z axis.

SOLUTION

We choose as the differential element of mass the thin slab shown; thus

$$dm = \rho bc \ dx$$

Referring to Sec. 9.13, we find that the moment of inertia of the element with respect to the z^\prime axis is

$$dI_{z'} = \frac{1}{12}b^2 dm$$

Applying the parallel-axis theorem, we obtain the mass moment of inertia of the slab with respect to the z axis.

$$dI_{z} = dI_{z'} + x^{2} dm = \frac{1}{12}b^{2} dm + x^{2} dm = (\frac{1}{12}b^{2} + x^{2})\rho bc dx$$

Integrating from x = 0 to x = a, we obtain

$$I_z = \int dI_z = \int_0^a \left(\frac{1}{12}b^2 + x^2\right)\rho bc \, dx = \rho abc\left(\frac{1}{12}b^2 + \frac{1}{3}a^2\right)$$

Since the total mass of the prism is $m = \rho a b c$, we can write

$$I_{z} = m(\frac{1}{12}b^{2} + \frac{1}{3}a^{2}) \qquad I_{z} = \frac{1}{12}m(4a^{2} + b^{2}) \quad \blacktriangleleft$$

We note that if the prism is thin, b is small compared to a, and the expression for I_z reduces to $\frac{1}{3}ma^2$, which is the result obtained in Sample Prob. 9.9 when L = a.



SAMPLE PROBLEM 9.11

Determine the moment of inertia of a right circular cone with respect to (a) its longitudinal axis, (b) an axis through the apex of the cone and perpendicular to its longitudinal axis, (c) an axis through the centroid of the cone and perpendicular to its longitudinal axis.



SOLUTION

We choose the differential element of mass shown.

$$r = a \frac{x}{h}$$
 $dm =
ho \pi r^2 dx =
ho \pi \frac{a^2}{h^2} x^2 dx$

a. Moment of Inertia I_x . Using the expression derived in Sec. 9.13 for a thin disk, we compute the mass moment of inertia of the differential element with respect to the x axis.

$$dI_x = \frac{1}{2}r^2 dm = \frac{1}{2} \left(a\frac{x}{h} \right)^2 \left(\rho \pi \frac{a^2}{h^2} x^2 dx \right) = \frac{1}{2} \rho \pi \frac{a^4}{h^4} x^4 dx$$

Integrating from x = 0 to x = h, we obtain

$$I_x = \int dI_x = \int_0^h \frac{1}{2} \rho \pi \frac{a^4}{h^4} x^4 \, dx = \frac{1}{2} \rho \pi \frac{a^4}{h^4} \frac{h^5}{5} = \frac{1}{10} \rho \pi a^4 h$$

Since the total mass of the cone is $m = \frac{1}{3}\rho\pi a^2 h$, we can write

$$I_{x} = \frac{1}{10}\rho\pi a^{4}h = \frac{3}{10}a^{2}(\frac{1}{3}\rho\pi a^{2}h) = \frac{3}{10}ma^{2} \qquad I_{x} = \frac{3}{10}ma^{2} \quad \blacktriangleleft$$

b. Moment of Inertia l_y . The same differential element is used. Applying the parallel-axis theorem and using the expression derived in Sec. 9.13 for a thin disk, we write

$$dI_y = dI_{y'} + x^2 dm = \frac{1}{4}r^2 dm + x^2 dm = (\frac{1}{4}r^2 + x^2) dm$$

Substituting the expressions for r and dm into the equation, we obtain

$$dI_y = \left(\frac{1}{4}\frac{a^2}{h^2}x^2 + x^2\right) \left(\rho\pi\frac{a^2}{h^2}x^2\,dx\right) = \rho\pi\frac{a^2}{h^2}\left(\frac{a^2}{4h^2} + 1\right)x^4\,dx$$
$$I_y = \int dI_y = \int_0^h \rho\pi\frac{a^2}{h^2}\left(\frac{a^2}{4h^2} + 1\right)x^4\,dx = \rho\pi\frac{a^2}{h^2}\left(\frac{a^2}{4h^2} + 1\right)\frac{h^5}{5}$$

Introducing the total mass of the cone m, we rewrite I_y as follows:

$$I_y = \frac{3}{5}(\frac{1}{4}a^2 + h^2)\frac{1}{3}\rho\pi a^2h \qquad I_y = \frac{3}{5}m(\frac{1}{4}a^2 + h^2) \quad \blacktriangleleft$$

c. Moment of Inertia $I_{y''}$. We apply the parallel-axis theorem and write $I_y = \overline{I}_{y''} + m\overline{x}^2$





CHAPTER Systems of Particles

Chapter 14 Systems of Particles

14.1 Introduction

- 14.2 Application of Newton's Laws to the Motion of a System of Particles. Effective Forces
- 14.3 Linear and Angular Momentum of a System of Particles
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14.1 INTRODUCTION

In this chapter you will study the motion of *systems of particles*, i.e., the motion of a large number of particles considered together. The first part of the chapter is devoted to systems consisting of well-defined particles; the second part considers the motion of variable systems, i.e., systems which are continually gaining or losing particles, or doing both at the same time.

In Sec. 14.2, Newton's second law will first be applied to each particle of the system. Defining the *effective force* of a particle as the product $m_i \mathbf{a}_i$ of its mass m_i and its acceleration \mathbf{a}_i , we will show that the *external forces* acting on the various particles form a system equipollent to the system of the effective forces, i.e., both systems have the same resultant and the same moment resultant about any given point. In Sec. 14.3, it will be further shown that the resultant and moment resultant of the external forces are equal, respectively, to the rate of change of the total linear momentum and of the total angular momentum of the particles of the system.

In Sec. 14.4, the *mass center* of a system of particles is defined and the motion of that point is described, while in Sec. 14.5 the motion of the particles about their mass center is analyzed. The conditions under which the linear momentum and the angular momentum of a system of particles are conserved are discussed in Sec. 14.6, and the results obtained in that section are applied to the solution of various problems.

Sections 14.7 and 14.8 deal with the application of the workenergy principle to a system of particles, and Sec. 14.9 with the application of the impulse-momentum principle. These sections also contain a number of problems of practical interest.

It should be noted that while the derivations given in the first part of this chapter are carried out for a system of independent particles, they remain valid when the particles of the system are rigidly connected, i.e., when they form a rigid body. In fact, the results obtained here will form the foundation of our discussion of the kinetics of rigid bodies in Chaps. 16 through 18.

The second part of this chapter is devoted to the study of variable systems of particles. In Sec. 14.11 you will consider steady streams of particles, such as a stream of water diverted by a fixed vane, or the flow of air through a jet engine, and learn to determine the force exerted by the stream on the vane and the thrust developed by the engine. Finally, in Sec. 14.12, you will learn how to analyze systems which gain mass by continually absorbing particles or lose mass by continually expelling particles. Among the various practical applications of this analysis will be the determination of the thrust developed by a rocket engine.

14.2 APPLICATION OF NEWTON'S LAWS TO THE MOTION OF A SYSTEM OF PARTICLES. EFFECTIVE FORCES

In order to derive the equations of motion for a system of n particles, let us begin by writing Newton's second law for each individual particle of the system. Consider the particle P_i , where $1 \le i \le n$. Let

 m_i be the mass of P_i and \mathbf{a}_i its acceleration with respect to the newtonian frame of reference *Oxyz*. The force exerted on P_i by another particle P_j of the system (Fig. 14.1), called an *internal force*, will be denoted by \mathbf{f}_{ij} . The resultant of the internal forces exerted on P_i by

all the other particles of the system is thus $\sum_{j=1}^{j} \mathbf{f}_{ij}$ (where \mathbf{f}_{ii} has no

meaning and is assumed to be equal to zero). Denoting, on the other hand, by \mathbf{F}_i the resultant of all the *external forces* acting on P_i , we write Newton's second law for the particle P_i as follows:

$$\mathbf{F}_i + \sum_{j=1}^n \mathbf{f}_{ij} = m_i \mathbf{a}_i \tag{14.1}$$

Denoting by \mathbf{r}_i the position vector of P_i and taking the moments about O of the various terms in Eq. (14.1), we also write

$$\mathbf{r}_i \times \mathbf{F}_i + \sum_{j=1}^n (\mathbf{r}_i \times \mathbf{f}_{ij}) = \mathbf{r}_i \times m_i \mathbf{a}_i$$
(14.2)

Repeating this procedure for each particle P_i of the system, we obtain n equations of the type (14.1) and n equations of the type (14.2), where i takes successively the values $1, 2, \ldots, n$. The vectors $m_i \mathbf{a}_i$ are referred to as the *effective forces* of the particles.[†] Thus the equations obtained express the fact that the external forces \mathbf{F}_i and the internal forces \mathbf{f}_{ij} acting on the various particles form a system equivalent to the system of the effective forces $m_i \mathbf{a}_i$ (i.e., one system may be replaced by the other) (Fig. 14.2).



Before proceeding further with our derivation, let us examine the internal forces \mathbf{f}_{ij} . We note that these forces occur in pairs \mathbf{f}_{ij} , \mathbf{f}_{ji} , where \mathbf{f}_{ij} represents the force exerted by the particle P_j on the particle P_i and \mathbf{f}_{ji} represents the force exerted by P_i on P_j (Fig. 14.2). Now, according to Newton's third law (Sec. 6.1), as extended by Newton's law of gravitation to particles acting at a distance (Sec. 12.10), the forces \mathbf{f}_{ij} and \mathbf{f}_{ji} are equal and opposite and have the same line of action. Their sum is therefore $\mathbf{f}_{ij} + \mathbf{f}_{ji} = 0$, and the sum of their moments about O is

 $\mathbf{r}_i \times \mathbf{f}_{ij} + \mathbf{r}_j \times \mathbf{f}_{ji} = \mathbf{r}_i \times (\mathbf{f}_{ij} + \mathbf{f}_{ji}) + (\mathbf{r}_j - \mathbf{r}_i) \times \mathbf{f}_{ji} = 0$

†Since these vectors represent the resultants of the forces acting on the various particles of the system, they can truly be considered as forces.





since the vectors $\mathbf{r}_j - \mathbf{r}_i$ and \mathbf{f}_{ji} in the last term are collinear. Adding all the internal forces of the system and summing their moments about O, we obtain the equations

$$\sum_{i=1}^{n} \sum_{j=1}^{n} \mathbf{f}_{ij} = 0 \qquad \sum_{i=1}^{n} \sum_{j=1}^{n} (\mathbf{r}_{i} \times \mathbf{f}_{ij}) = 0$$
(14.3)

which express the fact that the resultant and the moment resultant of the internal forces of the system are zero.

Returning now to the *n* equations (14.1), where i = 1, 2, ..., n, we sum their left-hand members and sum their right-hand members. Taking into account the first of Eqs. (14.3), we obtain

$$\sum_{i=1}^{n} \mathbf{F}_{i} = \sum_{i=1}^{n} m_{i} \mathbf{a}_{i}$$
(14.4)

Proceeding similarly with Eqs. (14.2) and taking into account the second of Eqs. (14.3), we have

$$\sum_{i=1}^{n} (\mathbf{r}_{i} \times \mathbf{F}_{i}) = \sum_{i=1}^{n} (\mathbf{r}_{i} \times m_{i} \mathbf{a}_{i})$$
(14.5)

Equations (14.4) and (14.5) express the fact that the system of the external forces \mathbf{F}_i and the system of the effective forces $m_i \mathbf{a}_i$ have the same resultant and the same moment resultant. Referring to the definition given in Sec. 3.19 for two equipollent systems of vectors, we can therefore state that the system of the external forces acting on the particles and the system of the effective forces of the particles are equipollent[†] (Fig. 14.3).



[†]The result just obtained is often referred to as *d'Alembert's principle*, after the French mathematician Jean le Rond d'Alembert (1717–1783). However, d'Alembert's original statement refers to the motion of a system of connected bodies, with \mathbf{f}_{ij} representing constraint forces which if applied by themselves will not cause the system to move. Since, as it will now be shown, this is in general not the case for the internal forces acting on a system of free particles, the consideration of d'Alembert's principle will be postponed until the motion of rigid bodies is considered (Chap. 16).

Equations (14.3) express the fact that the system of the internal forces \mathbf{f}_{ij} is equipollent to zero. Note, however, that it does *not* follow that the internal forces have no effect on the particles under consideration. Indeed, the gravitational forces that the sun and the planets exert on one another are internal to the solar system and equipollent to zero. Yet these forces are alone responsible for the motion of the planets about the sun.

Similarly, it does not follow from Eqs. (14.4) and (14.5) that two systems of external forces which have the same resultant and the same moment resultant will have the same effect on a given system of particles. Clearly, the systems shown in Figs. 14.4a and 14.4b have



the same resultant and the same moment resultant; yet the first system accelerates particle A and leaves particle B unaffected, while the second accelerates B and does not affect A. It is important to recall that when we stated in Sec. 3.19 that two equipollent systems of forces acting on a rigid body are also equivalent, we specifically noted that this property could *not* be extended to a system of forces acting on a set of independent particles such as those considered in this chapter.

In order to avoid any confusion, blue equals signs are used to connect equipollent systems of vectors, such as those shown in Figs. 14.3 and 14.4. These signs indicate that the two systems of vectors have the same resultant and the same moment resultant. Red equals signs will continue to be used to indicate that two systems of vectors are equivalent, i.e., that one system can actually be replaced by the other (Fig. 14.2).

14.3 LINEAR AND ANGULAR MOMENTUM OF A SYSTEM OF PARTICLES

Equations (14.4) and (14.5), obtained in the preceding section for the motion of a system of particles, can be expressed in a more condensed form if we introduce the linear and the angular momentum of the system of particles. Defining the linear momentum \mathbf{L} of the system of particles as the sum of the linear momenta of the various particles of the system (Sec. 12.3), we write

$$\mathbf{L} = \sum_{i=1}^{n} m_i \mathbf{v}_i \tag{14.6}$$

Defining the angular momentum \mathbf{H}_O about O of the system of particles in a similar way (Sec. 12.7), we have

$$\mathbf{H}_{O} = \sum_{i=1}^{n} \left(\mathbf{r}_{i} \times m_{i} \mathbf{v}_{i} \right)$$
(14.7)

Differentiating both members of Eqs. (14.6) and (14.7) with respect to t, we write

$$\dot{\mathbf{L}} = \sum_{i=1}^{n} m_i \dot{\mathbf{v}}_i = \sum_{i=1}^{n} m_i \mathbf{a}_i$$
(14.8)

and

$$\dot{\mathbf{H}}_{O} = \sum_{i=1}^{n} (\dot{\mathbf{r}}_{i} \times m_{i} \mathbf{v}_{i}) + \sum_{i=1}^{n} (\mathbf{r}_{i} \times m_{i} \dot{\mathbf{v}}_{i})$$
$$= \sum_{i=1}^{n} (\mathbf{v}_{i} \times m_{i} \mathbf{v}_{i}) + \sum_{i=1}^{n} (\mathbf{r}_{i} \times m_{i} \mathbf{a}_{i})$$

which reduces to

$$\dot{\mathbf{H}}_{O} = \sum_{i=1}^{n} \left(\mathbf{r}_{i} \times m_{i} \mathbf{a}_{i} \right)$$
(14.9)

since the vectors \mathbf{v}_i and $m_i \mathbf{v}_i$ are collinear.

We observe that the right-hand members of Eqs. (14.8) and (14.9) are respectively identical with the right-hand members of Eqs. (14.4) and (14.5). It follows that the left-hand members of these equations are respectively equal. Recalling that the left-hand member of Eq. (14.5) represents the sum of the moments \mathbf{M}_O about O of the external forces acting on the particles of the system, and omitting the subscript i from the sums, we write

$$\Sigma \mathbf{F} = \dot{\mathbf{L}}$$
(14.10)
$$\Sigma \mathbf{M}_O = \dot{\mathbf{H}}_O$$
(14.11)

These equations express that the resultant and the moment resultant about the fixed point O of the external forces are respectively equal to the rates of change of the linear momentum and of the angular momentum about O of the system of particles.

14.4 MOTION OF THE MASS CENTER OF A SYSTEM OF PARTICLES

Equation (14.10) may be written in an alternative form if the *mass* center of the system of particles is considered. The mass center of the system is the point G defined by the position vector $\mathbf{\bar{r}}$, which

satisfies the relation

$$m\overline{\mathbf{r}} = \sum_{i=1}^{n} m_i \mathbf{r}_i$$
(14.12)

where *m* represents the total mass $\sum_{i=1}^{n} m_i$ of the particles. Resolving the position vectors $\overline{\mathbf{r}}$ and \mathbf{r}_i into rectangular components, we obtain the following three scalar equations, which can be used to determine the coordinates $\overline{x}, \overline{y}, \overline{z}$ of the mass center:

$$m\overline{x} = \sum_{i=1}^{n} m_i x_i \qquad m\overline{y} = \sum_{i=1}^{n} m_i y_i \qquad m\overline{z} = \sum_{i=1}^{n} m_i z_i \qquad (14.12')$$

Since $m_i g$ represents the weight of the particle P_i , and mg the total weight of the particles, G is also the center of gravity of the system of particles. However, in order to avoid any confusion, G will be referred to as the *mass center* of the system of particles when properties associated with the *mass* of the particles are being discussed, and as the *center of gravity* of the system when properties associated with the *weight* of the particles are being considered. Particles located outside the gravitational field of the earth, for example, have a mass but no weight. We can then properly refer to their mass center, but obviously not to their center of gravity.[†]

Differentiating both members of Eq. (14.12) with respect to t, we write

$$m\dot{\overline{\mathbf{r}}} = \sum_{i=1}^{n} m_i \dot{\mathbf{r}}_i$$

or

$$m\overline{\mathbf{v}} = \sum_{i=1}^{n} m_i \mathbf{v}_i \tag{14.13}$$

where $\overline{\mathbf{v}}$ represents the velocity of the mass center *G* of the system of particles. But the right-hand member of Eq. (14.13) is, by definition, the linear momentum **L** of the system (Sec. 14.3). We therefore have

$$\mathbf{L} = m\overline{\mathbf{v}} \tag{14.14}$$

and, differentiating both members with respect to t,

$$\dot{\mathbf{L}} = m\overline{\mathbf{a}} \tag{14.15}$$

[†]It may also be pointed out that the mass center and the center of gravity of a system of particles do not exactly coincide, since the weights of the particles are directed toward the center of the earth and thus do not truly form a system of parallel forces.

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where $\overline{\mathbf{a}}$ represents the acceleration of the mass center G. Substituting for $\mathbf{\dot{L}}$ from (14.15) into (14.10), we write the equation

 $\Sigma \mathbf{F} = m \overline{\mathbf{a}} \tag{14.16}$

which defines the motion of the mass center G of the system of particles.

We note that Eq. (14.16) is identical with the equation we would obtain for a particle of mass m equal to the total mass of the particles of the system, acted upon by all the external forces. We therefore state that the mass center of a system of particles moves as if the entire mass of the system and all the external forces were concentrated at that point.

This principle is best illustrated by the motion of an exploding shell. We know that if air resistance is neglected, it can be assumed that a shell will travel along a parabolic path. After the shell has exploded, the mass center G of the fragments of shell will continue to travel along the same path. Indeed, point G must move as if the mass and the weight of all fragments were concentrated at G; it must, therefore, move as if the shell had not exploded.

It should be noted that the preceding derivation does not involve the moments of the external forces. Therefore, *it would be wrong to assume* that the external forces are equipollent to a vector $m\overline{\mathbf{a}}$ attached at the mass center *G*. This is not in general the case since, as you will see in the next section, the sum of the moments about *G* of the external forces is not in general equal to zero.

14.5 ANGULAR MOMENTUM OF A SYSTEM OF PARTICLES ABOUT ITS MASS CENTER

In some applications (for example, in the analysis of the motion of a rigid body) it is convenient to consider the motion of the particles of the system with respect to a centroidal frame of reference Gx'y'z'which translates with respect to the newtonian frame of reference Oxyz (Fig. 14.5). Although a centroidal frame is not, in general, a newtonian frame of reference, it will be seen that the fundamental relation (14.11) holds when the frame Oxyz is replaced by Gx'y'z'.

Denoting, respectively, by \mathbf{r}'_i and \mathbf{v}'_i the position vector and the velocity of the particle P_i relative to the moving frame of reference Gx'y'z', we define the *angular momentum* \mathbf{H}'_G of the system of particles *about the mass center* G as follows:

$$\mathbf{H}_{G}^{\prime} = \sum_{i=1}^{n} \left(\mathbf{r}_{i}^{\prime} \times m_{i} \mathbf{v}_{i}^{\prime} \right)$$
(14.17)

We now differentiate both members of Eq. (14.17) with respect to *t*. This operation is similar to that performed in Sec. 14.3 on Eq. (14.7), and so we write immediately

$$\dot{\mathbf{H}}_{G}^{\prime} = \sum_{i=1}^{n} \left(\mathbf{r}_{i}^{\prime} \times m_{i} \mathbf{a}_{i}^{\prime} \right)$$
(14.18)





where \mathbf{a}'_i denotes the acceleration of P_i relative to the moving frame of reference. Referring to Sec. 11.12, we write

$$\mathbf{a}_i = \overline{\mathbf{a}} + \mathbf{a}'_i$$

where \mathbf{a}_i and $\overline{\mathbf{a}}$ denote, respectively, the accelerations of P_i and G relative to the frame *Oxyz*. Solving for \mathbf{a}'_i and substituting into (14.18), we have

$$\dot{\mathbf{H}}_{G}^{\prime} = \sum_{i=1}^{n} \left(\mathbf{r}_{i}^{\prime} \times m_{i} \mathbf{a}_{i} \right) - \left(\sum_{i=1}^{n} m_{i} \mathbf{r}_{i}^{\prime} \right) \times \overline{\mathbf{a}}$$
(14.19)

But, by (14.12), the second sum in Eq. (14.19) is equal to $m\overline{\mathbf{r}}'$ and thus to zero, since the position vector $\overline{\mathbf{r}}'$ of G relative to the frame Gx'y'z' is clearly zero. On the other hand, since \mathbf{a}_i represents the acceleration of P_i relative to a newtonian frame, we can use Eq. (14.1) and replace $m_i \mathbf{a}_i$ by the sum of the internal forces \mathbf{f}_{ij} and of the resultant \mathbf{F}_i of the external forces acting on P_i . But a reasoning similar to that used in Sec. 14.2 shows that the moment resultant about G of the internal forces \mathbf{f}_{ij} of the entire system is zero. The first sum in Eq. (14.19) therefore reduces to the moment resultant about G of the external forces acting on the particles of the system, and we write

$$\Sigma \mathbf{M}_G = \dot{\mathbf{H}}'_G \tag{14.20}$$

which expresses that the moment resultant about G of the external forces is equal to the rate of change of the angular momentum about G of the system of particles.

It should be noted that in Eq. (14.17) we defined the angular momentum \mathbf{H}'_G as the sum of the moments about G of the momenta of the particles $m_i \mathbf{v}'_i$ in their motion relative to the centroidal frame of reference Gx'y'z'. We may sometimes want to compute the sum \mathbf{H}_G of the moments about G of the momenta of the particles $m_i \mathbf{v}_i$ in their absolute motion, i.e., in their motion as observed from the newtonian frame of reference Oxyz (Fig. 14.6):

$$\mathbf{H}_{G} = \sum_{i=1}^{n} \left(\mathbf{r}'_{i} \times m_{i} \mathbf{v}_{i} \right)$$
(14.21)

Remarkably, the angular momenta \mathbf{H}'_{G} and \mathbf{H}_{G} are identically equal. This can be verified by referring to Sec. 11.12 and writing

$$\mathbf{v}_i = \overline{\mathbf{v}} + \mathbf{v}'_i \tag{14.22}$$

Substituting for \mathbf{v}_i from (14.22) into Eq. (14.21), we have

$$\mathbf{H}_{G} = \left(\sum_{i=1}^{n} m_{i} \mathbf{r}_{i}^{\prime}\right) \times \mathbf{\bar{v}} + \sum_{i=1}^{n} \left(\mathbf{r}_{i}^{\prime} \times m_{i} \mathbf{v}_{i}^{\prime}\right)$$

But, as observed earlier, the first sum is equal to zero. Thus \mathbf{H}_G reduces to the second sum, which, by definition, is equal to \mathbf{H}'_G .[†]

[†]Note that this property is peculiar to the centroidal frame Gx'y'z' and does not, in general, hold for other frames of reference (see Prob. 14.29).





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Taking advantage of the property we have just established, we simplify our notation by dropping the prime (') from Eq. (14.20) and writing

$$\Sigma \mathbf{M}_G = \dot{\mathbf{H}}_G \tag{14.23}$$

where it is understood that the angular momentum \mathbf{H}_G can be computed by forming the moments about G of the momenta of the particles in their motion with respect to either the newtonian frame Oxyz or the centroidal frame Gx'y'z':

$$\mathbf{H}_{G} = \sum_{i=1}^{n} \left(\mathbf{r}'_{i} \times m_{i} \mathbf{v}_{i} \right) = \sum_{i=1}^{n} \left(\mathbf{r}'_{i} \times m_{i} \mathbf{v}'_{i} \right)$$
(14.24)



Chapter 16 Plane Motion of Rigid Bodies: Forces and Accelerations

- 16.1 Introduction
- 16.2 Equations of Motion for a Rigid Body
- 16.3 Angular Momentum of a Rigid Body in Plane Motion
- 16.4 Plane Motion of a Rigid Body. D'Alembert's Principle
- 16.5 A Remark on the Axioms of the Mechanics of Rigid Bodies
- 16.6 Solution of Problems Involving the Motion of a Rigid Body
- 16.7 Systems of Rigid Bodies
- 16.8 Constrained Plane Motion

16.1 INTRODUCTION

In this chapter and in Chaps. 17 and 18, you will study the *kinetics* of rigid bodies, i.e., the relations existing between the forces acting on a rigid body, the shape and mass of the body, and the motion produced. In Chaps. 12 and 13, you studied similar relations, assuming then that the body could be considered as a particle, i.e., that its mass could be concentrated in one point and that all forces acted at that point. The shape of the body, as well as the exact location of the points of application of the forces, will now be taken into account. You will also be concerned not only with the motion of the body as a whole but also with the motion of the body about its mass center.

Our approach will be to consider rigid bodies as made of large numbers of particles and to use the results obtained in Chap. 14 for the motion of systems of particles. Specifically, two equations from Chap. 14 will be used: Eq. (14.16), $\Sigma \mathbf{F} = m \mathbf{\bar{a}}$, which relates the resultant of the external forces and the acceleration of the mass center *G* of the system of particles, and Eq. (14.23), $\Sigma \mathbf{M}_G = \mathbf{\dot{H}}_G$, which relates the moment resultant of the external forces and the angular momentum of the system of particles about *G*.

Except for Sec. 16.2, which applies to the most general case of the motion of a rigid body, the results derived in this chapter will be limited in two ways: (1) They will be restricted to the *plane motion* of rigid bodies, i.e., to a motion in which each particle of the body remains at a constant distance from a fixed reference plane. (2) The rigid bodies considered will consist only of plane slabs and of bodies which are symmetrical with respect to the reference plane.† The study of the plane motion of nonsymmetrical three-dimensional bodies and, more generally, the motion of rigid bodies in three-dimensional space will be postponed until Chap. 18.

In Sec. 16.3, we define the angular momentum of a rigid body in plane motion and show that the rate of change of the angular momentum $\dot{\mathbf{H}}_{G}$ about the mass center is equal to the product $\bar{I}\boldsymbol{\alpha}$ of the centroidal mass moment of inertia \bar{I} and the angular acceleration $\boldsymbol{\alpha}$ of the body. D'Alembert's principle, introduced in Sec. 16.4, is used to prove that the external forces acting on a rigid body are equivalent to a vector $m\bar{\mathbf{a}}$ attached at the mass center and a couple of moment $\bar{I}\boldsymbol{\alpha}$.

In Sec. 16.5, we derive the principle of transmissibility using only the parallelogram law and Newton's laws of motion, allowing us to remove this principle from the list of axioms (Sec. 1.2) required for the study of the statics and dynamics of rigid bodies.

Free-body-diagram equations are introduced in Sec. 16.6 and will be used in the solution of all problems involving the plane motion of rigid bodies.

After considering the plane motion of connected rigid bodies in Sec. 16.7, you will be prepared to solve a variety of problems involving the translation, centroidal rotation, and unconstrained motion of rigid bodies. In Sec. 16.8 and in the remaining part of the chapter, the solution of problems involving noncentroidal rotation, rolling motion, and other partially constrained plane motions of rigid bodies will be considered.

 \dagger Or, more generally, bodies which have a principal centroidal axis of inertia perpendicular to the reference plane.

16.2 EQUATIONS OF MOTION FOR A RIGID BODY

Consider a rigid body acted upon by several external forces \mathbf{F}_1 , \mathbf{F}_2 , \mathbf{F}_3 , ... (Fig. 16.1). We can assume that the body is made of a large number n of particles of mass Δm_i (i = 1, 2, ..., n) and apply the results obtained in Chap. 14 for a system of particles (Fig. 16.2). Considering first the motion of the mass center G of the body with respect to the newtonian frame of reference Oxyz, we recall Eq. (14.16) and write

$$\Sigma \mathbf{F} = m \overline{\mathbf{a}} \tag{16.1}$$

where *m* is the mass of the body and $\overline{\mathbf{a}}$ is the acceleration of the mass center *G*. Turning now to the motion of the body relative to the centroidal frame of reference Gx'y'z', we recall Eq. (14.23) and write

$$\Sigma \mathbf{M}_G = \dot{\mathbf{H}}_G \tag{16.2}$$

where $\dot{\mathbf{H}}_{G}$ represents the rate of change of \mathbf{H}_{G} , the angular momentum about G of the system of particles forming the rigid body. In the following, \mathbf{H}_{G} will simply be referred to as the angular momentum of the rigid body about its mass center G. Together Eqs. (16.1) and (16.2) express that the system of the external forces is equipollent to the system consisting of the vector $m\mathbf{\bar{a}}$ attached at G and the couple of moment $\dot{\mathbf{H}}_{G}$ (Fig. 16.3).†



Equations (16.1) and (16.2) apply in the most general case of the motion of a rigid body. In the rest of this chapter, however, our analysis will be limited to the *plane motion* of rigid bodies, i.e., to a motion in which each particle remains at a constant distance from a fixed reference plane, and it will be assumed that the rigid bodies considered consist only of plane slabs and of bodies which are symmetrical with respect to the reference plane. Further study of the plane motion of nonsymmetrical three-dimensional bodies and of the motion of rigid bodies in three-dimensional space will be postponed until Chap. 18.

†Since the systems involved act on a rigid body, we could conclude at this point, by referring to Sec. 3.19, that the two systems are *equivalent* as well as equipollent and use red rather than blue equals signs in Fig. 16.3. However, by postponing this conclusion, we will be able to arrive at it independently (Secs. 16.4 and 18.5), thereby eliminating the necessity of including the principle of transmissibility among the axioms of mechanics (Sec. 16.5).

Photo 16.1 The system of external forces acting on the man and wakeboard includes the weights, the tension in the tow rope, and the forces exerted by the water and the air.



16.2 Equations of Motion for a Rigid Body **1027**

 \mathbf{F}_3

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16.3 ANGULAR MOMENTUM OF A RIGID BODY IN PLANE MOTION

Consider a rigid slab in plane motion. Assuming that the slab is made of a large number n of particles P_i of mass Δm_i and recalling Eq. (14.24) of Sec. 14.5, we note that the angular momentum \mathbf{H}_G of the slab about its mass center G can be computed by taking the moments about G of the momenta of the particles of the slab in their motion with respect to either of the frames Oxy or Gx'y' (Fig. 16.4). Choosing the latter course, we write

$$\mathbf{H}_{G} = \sum_{i=1}^{n} \left(\mathbf{r}_{i}^{\prime} \times \mathbf{v}_{i}^{\prime} \,\Delta m_{i} \right) \tag{16.3}$$

where \mathbf{r}'_i and $\mathbf{v}'_i \Delta m_i$ denote, respectively, the position vector and the linear momentum of the particle P_i relative to the centroidal frame of reference Gx'y'. But since the particle belongs to the slab, we have $\mathbf{v}'_i = \boldsymbol{\omega} \times \mathbf{r}'_i$, where $\boldsymbol{\omega}$ is the angular velocity of the slab at the instant considered. We write

$$\mathbf{H}_{G} = \sum_{i=1}^{n} \left[\mathbf{r}_{i}^{\prime} \times (\boldsymbol{\omega} \times \mathbf{r}_{i}^{\prime}) \Delta m_{i} \right]$$

Referring to Fig. 16.4, we easily verify that the expression obtained represents a vector of the same direction as $\boldsymbol{\omega}$ (that is, perpendicular to the slab) and of magnitude equal to $\omega \Sigma r_i^{\prime 2} \Delta m_i$. Recalling that the sum $\Sigma r_i^{\prime 2} \Delta m_i$ represents the moment of inertia \bar{I} of the slab about a centroidal axis perpendicular to the slab, we conclude that the angular momentum \mathbf{H}_G of the slab about its mass center is

$$\mathbf{H}_G = \bar{I}\boldsymbol{\omega} \tag{16.4}$$

Differentiating both members of Eq. (16.4) we obtain

$$\dot{\mathbf{H}}_{G} = \overline{I}\dot{\boldsymbol{\omega}} = \overline{I}\boldsymbol{\alpha} \tag{16.5}$$

Thus the rate of change of the angular momentum of the slab is represented by a vector of the same direction as $\boldsymbol{\alpha}$ (that is, perpendicular to the slab) and of magnitude $\bar{I}\alpha$.

It should be kept in mind that the results obtained in this section have been derived for a rigid slab in plane motion. As you will see in Chap. 18, they remain valid in the case of the plane motion of rigid bodies which are symmetrical with respect to the reference plane.[†] However, they do not apply in the case of nonsymmetrical bodies or in the case of three-dimensional motion.



Photo 16.2 The hard disk and pick-up arms of a hard disk computer undergo centroidal rotation.

[†]Or, more generally, bodies which have a principal centroidal axis of inertia perpendicular to the reference plane.





SAMPLE PROBLEM 16.2

The thin plate ABCD of mass 8 kg is held in the position shown by the wire BH and two links AE and DF. Neglecting the mass of the links, determine immediately after wire BH has been cut (a) the acceleration of the plate, (b) the force in each link.

SOLUTION

Kinematics of Motion. After wire BH has been cut, we observe that corners A and D move along parallel circles of radius 150 mm centered, respectively, at E and F. The motion of the plate is thus a curvilinear translation; the particles forming the plate move along parallel circles of radius 150 mm.

At the instant wire BH is cut, the velocity of the plate is zero. Thus the acceleration $\overline{\mathbf{a}}$ of the mass center G of the plate is tangent to the circular path which will be described by G.

Equations of Motion. The external forces consist of the weight **W** and the forces \mathbf{F}_{AE} and \mathbf{F}_{DF} exerted by the links. Since the plate is in translation, the effective forces reduce to the vector $m\overline{\mathbf{a}}$ attached at *G* and directed along the *t* axis. A free-body-diagram equation is drawn to show that the system of the external forces is equivalent to the system of the effective forces.

a. Acceleration of the Plate.

$$W \cos 30^{\circ} = m\overline{a}$$

$$mg \cos 30^{\circ} = m\overline{a}$$

$$\overline{a} = g \cos 30^{\circ} = (9.81 \text{ m/s}^2) \cos 30^{\circ} \qquad (1)$$

$$\overline{a} = 8.50 \text{ m/s}^2 \swarrow 60^{\circ} \blacktriangleleft$$

b. Forces in Links AE and DF.

 $+\nabla \Sigma F_n = \Sigma (F_n)_{\text{eff}}; \qquad F_{AE} + F_{DF} - W \sin 30^\circ = 0$ (2) $+ \sum M_G = \Sigma (M_G)_{\text{eff}};$

$$\begin{array}{l} (F_{AE}\,\sin\,30^\circ)(250\,\,\mathrm{mm})\,-\,(F_{AE}\,\cos\,30^\circ)(100\,\,\mathrm{mm}) \\ +\,(F_{DF}\,\sin\,30^\circ)(250\,\,\mathrm{mm})\,+\,(F_{DF}\,\cos\,30^\circ)(100\,\,\mathrm{mm})\,=\,0 \\ 38.4F_{AE}\,+\,211.6F_{DF}\,=\,0 \\ F_{DF}\,=\,-0.1815F_{AE} \eqno(3) \end{array}$$

Substituting for F_{DF} from (3) into (2), we write

$$F_{AE} - 0.1815F_{AE} - W \sin 30^{\circ} = 0$$

$$F_{AE} = 0.6109W$$

$$F_{DF} = -0.1815(0.6109W) = -0.1109W$$

Noting that $W = mg = (8 \text{ kg})(9.81 \text{ m/s}^2) = 78.48 \text{ N}$, we have

$$F_{AE} = 0.6109(78.48 \text{ N}) \qquad F_{AE} = 47.9 \text{ N} T \blacktriangleleft$$

$$F_{DF} = -0.1109(78.48 \text{ N}) \qquad F_{DF} = 8.70 \text{ N} C \blacktriangleleft$$





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 r_B

SAMPLE PROBLEM 16.3

A pulley weighing 12 lb and having a radius of gyration of 8 in. is connected to two blocks as shown. Assuming no axle friction, determine the angular acceleration of the pulley and the acceleration of each block.

SOLUTION

Sense of Motion. Although an arbitrary sense of motion can be assumed (since no friction forces are involved) and later checked by the sign of the answer, we may prefer to determine the actual sense of rotation of the pulley first. The weight of block B required to maintain the equilibrium of the pulley when it is acted upon by the 5-lb block A is first determined. We write

$$+\gamma \Sigma M_G = 0$$
: $W_B(6 \text{ in.}) - (5 \text{ lb})(10 \text{ in.}) = 0$ $W_B = 8.33 \text{ lb}$

Since block B actually weighs 10 lb, the pulley will rotate counterclockwise.

Kinematics of Motion. Assuming α counterclockwise and noting that $a_A = r_A \alpha$ and $a_B = r_B \alpha$, we obtain

$$\mathbf{a}_A = \begin{pmatrix} \frac{10}{12} & \text{ft} \end{pmatrix} \alpha \uparrow \qquad \mathbf{a}_B = \begin{pmatrix} \frac{6}{12} & \text{ft} \end{pmatrix} \alpha \downarrow$$

Equations of Motion. A single system consisting of the pulley and the two blocks is considered. Forces external to this system consist of the weights of the pulley and the two blocks and of the reaction at *G*. (The forces exerted by the cables on the pulley and on the blocks are internal to the system considered and cancel out.) Since the motion of the pulley is a centroidal rotation and the motion of each block is a translation, the effective forces reduce to the couple $I\alpha$ and the two vectors $m\mathbf{a}_A$ and $m\mathbf{a}_B$. The centroidal moment of inertia of the pulley is

$$\bar{I} = m\bar{k}^2 = \frac{W}{g}\bar{k}^2 = \frac{12 \text{ lb}}{32.2 \text{ ft/s}^2} (\frac{8}{12} \text{ ft})^2 = 0.1656 \text{ lb} \cdot \text{ft} \cdot \text{s}^2$$

Since the system of the external forces is equipollent to the system of the effective forces, we write

$$\begin{split} + & \gamma \Sigma M_G = \Sigma (M_G)_{\text{eff}}: \\ & (10 \text{ lb})(\frac{6}{12} \text{ ft}) - (5 \text{ lb})(\frac{10}{12} \text{ ft}) = + \overline{I}\alpha + m_B a_B(\frac{6}{12} \text{ ft}) + m_A a_A(\frac{10}{12} \text{ ft}) \\ & (10)(\frac{6}{12}) - (5)(\frac{10}{12}) = 0.1656\alpha + \frac{10}{32.2}(\frac{6}{12}\alpha)(\frac{6}{12}) + \frac{5}{32.2}(\frac{10}{12}\alpha)(\frac{10}{12}) \\ & \alpha = + 2.374 \text{ rad/s}^2 \qquad \alpha = 2.37 \text{ rad/s}^2 \gamma \\ & a_A = r_A \alpha = (\frac{10}{12} \text{ ft})(2.374 \text{ rad/s}^2) \qquad \mathbf{a}_A = 1.978 \text{ ft/s}^2 \uparrow \\ & a_B = r_B \alpha = (\frac{6}{12} \text{ ft})(2.374 \text{ rad/s}^2) \qquad \mathbf{a}_B = 1.187 \text{ ft/s}^2 \downarrow \end{split}$$





T A G o^{-5 m}

SAMPLE PROBLEM 16.4

A cord is wrapped around a homogeneous disk of radius r = 0.5 m and mass m = 15 kg. If the cord is pulled upward with a force **T** of magnitude 180 N, determine (a) the acceleration of the center of the disk, (b) the angular acceleration of the disk, (c) the acceleration of the cord.

SOLUTION

Equations of Motion. We assume that the components $\overline{\mathbf{a}}_x$ and $\overline{\mathbf{a}}_y$ of the acceleration of the center are directed, respectively, to the right and upward and that the angular acceleration of the disk is counterclockwise. The external forces acting on the disk consist of the weight \mathbf{W} and the force \mathbf{T} exerted by the cord. This system is equivalent to the system of the effective forces, which consists of a vector of components $m\overline{\mathbf{a}}_x$ and $m\overline{\mathbf{a}}_y$ attached at G and a couple $I\alpha$. We write

$$\stackrel{+}{\rightarrow} \Sigma F_x = \Sigma(F_x)_{\text{eff}}: \qquad 0 = m\overline{a}_x \qquad \overline{\mathbf{a}}_x = \mathbf{0} \quad \blacktriangleleft \\ +\uparrow \Sigma F_y = \Sigma(F_y)_{\text{eff}}: \qquad T - W = m\overline{a}_y \\ \overline{a}_y = \frac{T - W}{m}$$

Since T=180 N, m=15 kg, and $W=(15~{\rm kg})(9.81~{\rm m/s}^2)=147.1$ N, we have

$$\overline{a}_{y} = \frac{180 \text{ N} - 147.1 \text{ N}}{15 \text{ kg}} = +2.19 \text{ m/s}^{2} \qquad \overline{\mathbf{a}}_{y} = 2.19 \text{ m/s}^{2} \uparrow \checkmark$$

$$+\gamma \Sigma M_{G} = \Sigma (M_{G})_{\text{eff}}: \qquad -Tr = \overline{I}\alpha$$

$$-Tr = (\frac{1}{2}mr^{2})\alpha$$

$$\alpha = -\frac{2T}{mr} = -\frac{2(180 \text{ N})}{(15 \text{ kg})(0.5 \text{ m})} = -48.0 \text{ rad/s}^{2}$$

$$\alpha = 48.0 \text{ rad/s}^{2} \downarrow \checkmark$$

Acceleration of Cord. Since the acceleration of the cord is equal to the tangential component of the acceleration of point *A* on the disk, we write

$$\mathbf{a}_{\text{cord}} = (\mathbf{a}_A)_t = \overline{\mathbf{a}} + (\mathbf{a}_{A/G})_t$$
$$= [2.19 \text{ m/s}^2 \uparrow] + [(0.5 \text{ m})(48 \text{ rad/s}^2)\uparrow]$$
$$\mathbf{a}_{\text{cord}} = 26.2 \text{ m/s}^2 \uparrow \checkmark$$





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SAMPLE PROBLEM 16.5

A uniform sphere of mass m and radius r is projected along a rough horizontal surface with a linear velocity $\overline{\mathbf{v}}_0$ and no angular velocity. Denoting by μ_k the coefficient of kinetic friction between the sphere and the floor, determine (a) the time t_1 at which the sphere will start rolling without sliding, (b) the linear velocity and angular velocity of the sphere at time t_1 .

SOLUTION

 $m\overline{a}$

Ĝ

Equations of Motion. The positive sense is chosen to the right for $\overline{\mathbf{a}}$ and clockwise for $\boldsymbol{\alpha}$. The external forces acting on the sphere consist of the weight \mathbf{W} , the normal reaction \mathbf{N} , and the friction force \mathbf{F} . Since the point of the sphere in contact with the surface is sliding to the right, the friction force \mathbf{F} is directed to the left. While the sphere is sliding, the magnitude of the friction force is $F = \mu_k N$. The effective forces consist of the vector $m\overline{\mathbf{a}}$ attached at G and the couple $\overline{I}\boldsymbol{\alpha}$. Expressing that the system of the external forces is equivalent to the system of the effective forces, we write

$$\begin{split} +\uparrow \Sigma F_y &= \Sigma (F_y)_{\text{eff}}: & N - W = 0\\ N &= W = mg & F = \mu_k N = \mu_k mg\\ \stackrel{+}{\longrightarrow} \Sigma F_x &= \Sigma (F_x)_{\text{eff}}: & -F = m\overline{a} & -\mu_k mg = m\overline{a} & \overline{a} = -\mu_k g\\ +\downarrow \Sigma M_G &= \Sigma (M_G)_{\text{eff}}: & Fr = \overline{I}\alpha \end{split}$$

Noting that $\overline{I} = \frac{2}{5}mr^2$ and substituting the value obtained for *F*, we write

$$(\mu_k mg)r = \frac{2}{5}mr^2\alpha$$
 $\alpha = \frac{5}{2}\frac{\mu_k g}{r}$

Kinematics of Motion. As long as the sphere both rotates and slides, its linear and angular motions are uniformly accelerated.

$$t = 0, \overline{v} = \overline{v}_0 \qquad \overline{v} = \overline{v}_0 + \overline{a}t = \overline{v}_0 - \mu_k gt \tag{1}$$

$$= 0, \omega_0 = 0 \qquad \omega = \omega_0 + \alpha t = 0 + \left(\frac{5}{2}\frac{\mu_k g}{r}\right)t \qquad (2)$$

The sphere will start rolling without sliding when the velocity \mathbf{v}_C of the point of contact *C* is zero. At that time, $t = t_1$, point *C* becomes the instantaneous center of rotation, and we have

$$\overline{v}_1 = r\omega_1 \tag{3}$$

Substituting in (3) the values obtained for \overline{v}_1 and ω_1 by making $t = t_1$ in (1) and (2), respectively, we write

$$\overline{v}_0 - \mu_k g t_1 = r \left(\frac{5}{2} \frac{\mu_k g}{r} t_1 \right) \qquad t_1 = \frac{2}{7} \frac{\overline{v}_0}{\mu_k g} \blacktriangleleft$$

Substituting for t_1 into (2), we have

t

$$\omega_1 = \frac{5}{2} \frac{\mu_k g}{r} t_1 = \frac{5}{2} \frac{\mu_k g}{r} \left(\frac{2}{7} \frac{\overline{v}_0}{\mu_k g}\right) \qquad \omega_1 = \frac{5}{7} \frac{\overline{v}_0}{r} \qquad \omega_1 = \frac{5}{7} \frac{\overline{v}_0}{r} \downarrow \blacktriangleleft$$
$$\overline{v}_1 = r \omega_1 = r \left(\frac{5}{7} \frac{\overline{v}_0}{r}\right) \qquad \overline{v}_1 = \frac{5}{7} \overline{v}_0 \qquad \mathbf{v}_1 = \frac{5}{7} \overline{v}_0 \rightarrow \blacktriangleleft$$





Fig. 16.11

16.8 CONSTRAINED PLANE MOTION

Most engineering applications deal with rigid bodies which are moving under given constraints. For example, cranks must rotate about a fixed axis, wheels must roll without sliding, and connecting rods must describe certain prescribed motions. In all such cases, definite relations exist between the components of the acceleration $\overline{\mathbf{a}}$ of the mass center G of the body considered and its angular acceleration $\boldsymbol{\alpha}$; the corresponding motion is said to be a *constrained motion*.

The solution of a problem involving a constrained plane motion calls first for a *kinematic analysis* of the problem. Consider, for example, a slender rod AB of length l and mass m whose extremities are connected to blocks of negligible mass which slide along horizontal and vertical frictionless tracks. The rod is pulled by a force \mathbf{P} applied at A (Fig. 16.11). We know from Sec. 15.8 that the acceleration $\overline{\mathbf{a}}$ of the mass center G of the rod can be determined at any given instant from the position of the rod, its angular velocity, and its angular acceleration at that instant. Suppose, for example, that the values of θ , ω , and α are known at a given instant and that we wish to determine the corresponding value of the force \mathbf{P} , as well as the reactions at A and B. We should first determine the components \overline{a}_x and \overline{a}_y of the acceleration of the mass center G by the method of Sec. 15.8. We next apply d'Alembert's principle (Fig. 16.12), using the expressions obtained for \overline{a}_x and \overline{a}_y . The unknown forces **P**, **N**_A, and **N**_B can then be determined by writing and solving the appropriate equations.



Suppose now that the applied force **P**, the angle θ , and the angular velocity ω of the rod are known at a given instant and that we wish to find the angular acceleration α of the rod and the components \overline{a}_x and \overline{a}_y of the acceleration of its mass center at that instant, as well as the reactions at A and B. The preliminary kinematic study of the problem will have for its object to express the components \overline{a}_x and \overline{a}_y of the acceleration of G in terms of the angular acceleration α of the rod. This will be done by first expressing the acceleration of a suitable reference point such as A in terms of the angular acceleration α . The components \overline{a}_x and \overline{a}_y of the acceleration of G can then be determined in terms of α , and the expressions obtained carried into Fig. 16.12. Three equations can then be derived in terms of α , N_A , and N_B and solved for the three unknowns (see Sample

Prob. 16.10). Note that the method of dynamic equilibrium can also be used to carry out the solution of the two types of problems we have considered (Fig. 16.13).

When a mechanism consists of *several moving parts*, the approach just described can be used with each part of the mechanism. The procedure required to determine the various unknowns is then similar to the procedure followed in the case of the equilibrium of a system of connected rigid bodies (Sec. 6.11).

Earlier, we analyzed two particular cases of constrained plane motion: the translation of a rigid body, in which the angular acceleration of the body is constrained to be zero, and the centroidal rotation, in which the acceleration $\overline{\mathbf{a}}$ of the mass center of the body is constrained to be zero. Two other particular cases of constrained plane motion are of special interest: the *noncentroidal rotation* of a rigid body and the *rolling motion* of a disk or wheel. These two cases can be analyzed by one of the general methods described above. However, in view of the range of their applications, they deserve a few special comments.

Noncentroidal Rotation. The motion of a rigid body constrained to rotate about a fixed axis which does not pass through its mass center is called *noncentroidal rotation*. The mass center *G* of the body moves along a circle of radius \overline{r} centered at the point *O*, where the axis of rotation intersects the plane of reference (Fig. 16.14). Denoting, respectively, by $\boldsymbol{\omega}$ and $\boldsymbol{\alpha}$ the angular velocity and the angular acceleration of the line *OG*, we obtain the following expressions for the tangential and normal components of the acceleration of *G*:

$$\overline{a}_t = \overline{r}\alpha \qquad \overline{a}_n = \overline{r}\omega^2 \tag{16.7}$$

Since line OG belongs to the body, its angular velocity $\boldsymbol{\omega}$ and its angular acceleration $\boldsymbol{\alpha}$ also represent the angular velocity and the angular acceleration of the body in its motion relative to G. Equations (16.7) define, therefore, the kinematic relation existing between the motion of the mass center G and the motion of the body about G. They should be used to eliminate \bar{a}_t and \bar{a}_n from the equations obtained by applying d'Alembert's principle (Fig. 16.15) or the method of dynamic equilibrium (Fig. 16.16).







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An interesting relation is obtained by equating the moments about the fixed point O of the forces and vectors shown, respectively, in parts a and b of Fig. 16.15. We write

$$+ \sum \Sigma M_O = \overline{I}\alpha + (m\overline{r}\alpha)\overline{r} = (\overline{I} + m\overline{r}^2)\alpha$$

But according to the parallel-axis theorem, we have $\overline{I} + m\overline{r}^2 = I_O$, where I_O denotes the moment of inertia of the rigid body about the fixed axis. We therefore write

$$\Sigma M_O = I_O \alpha \tag{16.8}$$

Although formula (16.8) expresses an important relation between the sum of the moments of the external forces about the fixed point O and the product $I_O\alpha$, it should be clearly understood that this formula does not mean that the system of the external forces is equivalent to a couple of moment $I_O\alpha$. The system of the effective forces, and thus the system of the external forces, reduces to a couple only when O coincides with G—that is, only when the rotation is centroidal (Sec. 16.4). In the more general case of noncentroidal rotation, the system of the external forces not reduce to a couple.

A particular case of noncentroidal rotation is of special interest the case of *uniform rotation*, in which the angular velocity $\boldsymbol{\omega}$ is constant. Since $\boldsymbol{\alpha}$ is zero, the inertia couple in Fig. 16.16 vanishes and the inertia vector reduces to its normal component. This component (also called *centrifugal force*) represents the tendency of the rigid body to break away from the axis of rotation.

Rolling Motion. Another important case of plane motion is the motion of a disk or wheel rolling on a plane surface. If the disk is constrained to roll without sliding, the acceleration $\overline{\mathbf{a}}$ of its mass center *G* and its angular acceleration $\boldsymbol{\alpha}$ are not independent. Assuming that the disk is balanced, so that its mass center and its geometric center coincide, we first write that the distance \overline{x} traveled by *G* during a rotation θ of the disk is $\overline{x} = r\theta$, where *r* is the radius of the disk. Differentiating this relation twice, we write

$$\overline{a} = r\alpha \tag{16.9}$$



Recalling that the system of the effective forces in plane motion reduces to a vector $m\overline{\mathbf{a}}$ and a couple $\overline{I}\boldsymbol{\alpha}$, we find that in the particular case of the rolling motion of a balanced disk, the effective forces reduce to a vector of magnitude $mr\alpha$ attached at G and to a couple of magnitude $\overline{I}\alpha$. We may thus express that the external forces are equivalent to the vector and couple shown in Fig. 16.17.

When a disk rolls without sliding, there is no relative motion between the point of the disk in contact with the ground and the ground itself. Thus as far as the computation of the friction force **F** is concerned, a rolling disk can be compared with a block at rest on a surface. The magnitude F of the friction force can have any value, as long as this value does not exceed the maximum value $F_m = \mu_s N$, where μ_s is the coefficient of static friction and N is the magnitude of the normal force. In the case of a rolling disk, the magnitude F of the friction force should therefore be determined independently of Nby solving the equation obtained from Fig. 16.17.

When *sliding is impending*, the friction force reaches its maximum value $F_m = \mu_s N$ and can be obtained from N.

When the disk *rotates and slides* at the same time, a relative motion exists between the point of the disk which is in contact with the ground and the ground itself, and the force of friction has the magnitude $F_k = \mu_k N$, where μ_k is the coefficient of kinetic friction. In this case, however, the motion of the mass center G of the disk and the rotation of the disk about G are independent, and \overline{a} is not equal to $r\alpha$.

These three different cases can be summarized as follows:

Rolling, no sliding:	$F \leq \mu_s N$	$\overline{a} = r\alpha$
Rolling, sliding impending:	$F = \mu_s N$	$\overline{a} = r\alpha$
Rotating and sliding:	$F = \mu_k N$	\overline{a} and α independent

When it is not known whether or not a disk slides, it should first be assumed that the disk rolls without sliding. If F is found smaller than or equal to $\mu_s N$, the assumption is proved correct. If F is found larger than $\mu_s N$, the assumption is incorrect and the problem should be started again, assuming rotating and sliding.

When a disk is *unbalanced*, i.e., when its mass center G does not coincide with its geometric center O, the relation (16.9) does not hold between \overline{a} and α . However, a similar relation holds between the magnitude a_O of the acceleration of the geometric center and the angular acceleration α of an unbalanced disk which rolls without sliding. We have

$$a_{O} = r\alpha \tag{16.10}$$

To determine \overline{a} in terms of the angular acceleration α and the angular velocity ω of the disk, we can use the relative-acceleration formula

$$\overline{\mathbf{a}} = \overline{\mathbf{a}}_G = \mathbf{a}_O + \mathbf{a}_{G/O} = \mathbf{a}_O + (\mathbf{a}_{G/O})_t + (\mathbf{a}_{G/O})_n$$
(16.11)

where the three component accelerations obtained have the directions indicated in Fig. 16.18 and the magnitudes $a_O = r\alpha$, $(a_{G/O})_t = (OG)\alpha$, and $(a_{G/O})_n = (OG)\omega^2$.

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Photo 16.4 As the ball hits the bowling alley, it first spins and slides, then rolls without sliding.



Fig. 16.18



SAMPLE PROBLEM 16.8

A sphere of radius r and weight W is released with no initial velocity on the incline and rolls without slipping. Determine (a) the minimum value of the coefficient of static friction compatible with the rolling motion, (b) the velocity of the center G of the sphere after the sphere has rolled 10 ft, (c) the velocity of G if the sphere were to move 10 ft down a frictionless 30° incline.

SOLUTION

a. Minimum μ_s for Rolling Motion. The external forces **W**, **N**, and **F** form a system equivalent to the system of effective forces represented by the vector $m\bar{\mathbf{a}}$ and the couple $\bar{I}\boldsymbol{\alpha}$. Since the sphere rolls without sliding, we have $\bar{a} = r\alpha$.

$$+ \sum M_C = \sum (M_C)_{\text{eff}}: \qquad (W \sin \theta)r = (m\overline{a})r + \overline{I}\alpha \\ (W \sin \theta)r = (mr\alpha)r + \overline{I}\alpha$$

Noting that m = W/g and $\overline{I} = \frac{2}{5}mr^2$, we write

$$(W \sin \theta)r = \left(\frac{W}{g}r\alpha\right)r + \frac{2}{5}\frac{W}{g}r^{2}\alpha \qquad \alpha = +\frac{5g\sin\theta}{7r}$$
$$\overline{a} = r\alpha = \frac{5g\sin\theta}{7} = \frac{5(32.2 \text{ ft/s}^{2})\sin 30^{\circ}}{7} = 11.50 \text{ ft/s}^{2}$$
$$+\Sigma F_{x} = \Sigma(F_{x})_{\text{eff}}: \qquad W \sin \theta - F = m\overline{a}$$
$$W \sin \theta - F = \frac{W}{g}\frac{5g\sin\theta}{7}$$
$$F = +\frac{2}{7}W \sin \theta = \frac{2}{7}W \sin 30^{\circ} \qquad \mathbf{F} = 0.143W \text{ Im} 30^{\circ}$$
$$+\mathscr{I}\Sigma F_{y} = \Sigma(F_{y})_{\text{eff}}: \qquad N - W \cos \theta = 0$$
$$N = W \cos \theta = 0.866W \qquad \mathbf{N} = 0.866W \text{ and } 0.86W \text{ an$$

x **b. Velocity of Rolling Sphere.** We have uniformly accelerated motion:

c. Velocity of Sliding Sphere. Assuming now no friction, we have F = 0 and obtain

$$+ \sum \Sigma M_G = \Sigma (M_G)_{\text{eff}}: \quad 0 = I\alpha \qquad \alpha = 0$$

+
$$\sum F_x = \Sigma (F_x)_{\text{eff}}: \qquad W \sin 30^\circ = m\overline{a} \qquad 0.50W = \frac{W}{g} \overline{a}$$

$$\overline{a} = +16.1 \text{ ft/s}^2 \qquad \overline{a} = 16.1 \text{ ft/s}^2 \implies 30^\circ$$

Substituting $\overline{a} = 16.1 \text{ ft/s}^2$ into the equations for uniformly accelerated motion, we obtain

$$\overline{v}^2 = \overline{v}_0^2 + 2\overline{a}(\overline{x} - \overline{x}_0) \qquad \overline{v}^2 = 0 + 2(16.1 \text{ ft/s}^2)(10 \text{ ft}) \\ \overline{v} = 17.94 \text{ ft/s} \qquad \overline{\mathbf{v}} = 17.94 \text{ ft/s} \ 30^\circ \ \checkmark$$

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SAMPLE PROBLEM 16.10

The extremities of a 4-ft rod weighing 50 lb can move freely and with no friction along two straight tracks as shown. If the rod is released with no velocity from the position shown, determine (a) the angular acceleration of the rod, (b) the reactions at A and B.

SOLUTION

Kinematics of Motion. Since the motion is constrained, the acceleration of *G* must be related to the angular acceleration $\boldsymbol{\alpha}$. To obtain this relation, we first determine the magnitude of the acceleration \mathbf{a}_A of point *A* in terms of $\boldsymbol{\alpha}$. Assuming that $\boldsymbol{\alpha}$ is directed counterclockwise and noting that $a_{B/A} = 4\alpha$, we write

$$\mathbf{a}_{B} = \mathbf{a}_{A} + \mathbf{a}_{B/A}$$
$$a_{B} \checkmark 45^{\circ}] = [a_{A} \rightarrow] + [4\alpha \not \simeq 60^{\circ}]$$

Noting that $\phi = 75^{\circ}$ and using the law of sines, we obtain

$$a_A = 5.46\alpha$$
 $a_B = 4.90\alpha$

The acceleration of G is now obtained by writing

 $\overline{\mathbf{a}} = \mathbf{a}_G = \mathbf{a}_A + \mathbf{a}_{G/A}$ $\overline{\mathbf{a}} = [5.46\alpha \rightarrow] + [2\alpha \not \sim 60^\circ]$

Resolving $\overline{\mathbf{a}}$ into *x* and *y* components, we obtain

 $\overline{a}_x = 5.46\alpha - 2\alpha \cos 60^\circ = 4.46\alpha \qquad \overline{\mathbf{a}}_x = 4.46\alpha \rightarrow \overline{a}_y = -2\alpha \sin 60^\circ = -1.732\alpha \qquad \overline{\mathbf{a}}_y = 1.732\alpha \downarrow$

Kinetics of Motion. We draw a free-body-diagram equation expressing that the system of the external forces is equivalent to the system of the effective forces represented by the vector of components $m\overline{\mathbf{a}}_x$ and $m\overline{\mathbf{a}}_y$ attached at G and the couple $I\alpha$. We compute the following magnitudes:

$$\overline{I} = \frac{1}{12}ml^2 = \frac{1}{12}\frac{50 \text{ lb}}{32.2 \text{ ft/s}^2}(4 \text{ ft})^2 = 2.07 \text{ lb} \cdot \text{ft} \cdot \text{s}^2 \qquad \overline{I}\alpha = 2.07\alpha$$

$$\overline{a}_x = \frac{50}{32.2}(4.46\alpha) = 6.93\alpha$$
 $m\overline{a}_y = -\frac{50}{32.2}(1.732\alpha) = -2.69\alpha$

Equations of Motion

m

E

$$\begin{array}{rcl} & + \gamma \Sigma M_E = \Sigma(M_E)_{\rm eff}: \\ & (50)(1.732) = (6.93\alpha)(4.46) + (2.69\alpha)(1.732) + 2.07\alpha \\ & \alpha = +2.30 \ {\rm rad/s}^2 & \alpha = 2.30 \ {\rm rad/s}^2 & \gamma \\ & \\ & ft \end{array} \xrightarrow{+} \Sigma F_x = \Sigma(F_x)_{\rm eff}: \qquad R_B \ {\rm sin} \ 45^\circ = (6.93)(2.30) = 15.94 \\ & R_B = 22.5 \ {\rm lb} \qquad R_B = 22.5 \ {\rm lb} \ \swarrow 45^\circ \\ & \\ & + \gamma \Sigma F_y = \Sigma(F_y)_{\rm eff}: \ R_A + R_B \ {\rm cos} \ 45^\circ - 50 = -(2.69)(2.30) \\ & R_A = -6.19 - 15.94 + 50 = 27.9 \ {\rm lb} \\ & \\ & \\ & R_A = 27.9 \ {\rm lb} \ \uparrow 4 \\ \end{array}$$



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